

706

Accession Number

12331768

Author

Cox P. Krips M. Neri R. Omont A. Gusten R. Menten KM. Wyrowski F. Weiss A. Beelen A. Gurwell MA. Dannerbauer H. Ivison RJ. Negrello M. Aretxaga I. Hughes DH. Auld R. Baes M. Blundell R. Buttiglione S. Cava A. Cooray A. Dariush A. Dunne L. Dye S. Eales SA. Frayer D. Fritz J. Gavazzi R. Hopwood R. Ibar E. Jarvis M. Maddox S. Michalowski M. Pascale E. Pohlen M. Rigby E. Smith DJB. Swinbank AM. Temi P. Valtchanov I. van der Werf P. de Zotti G.

Author Unabbreviated

Cox P.; Krips M.; Neri R.; Omont A.; Gusten R.; Menten K. M.; Wyrowski F.; Weiss A.; Beelen A.; Gurwell M. A.; Dannerbauer H.; Ivison R. J.; Negrello M.; Aretxaga I.; Hughes D. H.; Auld R.; Baes M.; Blundell R.; Buttiglione S.; Cava A.; Cooray A.; Dariush A.; Dunne L.; Dye S.; Eales S. A.; Frayer D.; Fritz J.; Gavazzi R.; Hopwood R.; Ibar E.; Jarvis M.; Maddox S.; Michalowski M.; Pascale E.; Pohlen M.; Rigby E.; Smith D. J. B.; Swinbank A. M.; Temi P.; Valtchanov I.; van der Werf P.; de Zotti G.

Author/Editor Affiliation

Cox P. Krips M. Neri R. : IRAM, 300 rue de la piscine, St. Martin d'Herès F-38406, France

Omont A. Gavazzi R. : CNRS, Université Pierre et Marie Curie, 98 bis boulevard Arago, Paris 75014, France

Gusten R. Menten KM. Wyrowski F. Weiss A. : MPIfR, Auf dem Hugel 69, Bonn 53121, Germany

Beelen A. : IAS, Université Paris-Sud, Bat. 121, Orsay F-91405, France

Gurwell MA. Blundell R. : Harvard-Smithsonian Center for Astrophysics, 60 Garden Street, Mail Stop 42, Cambridge, MA 02138, USA

Dannerbauer H. : CNRS, Université de Paris Diderot, Orme des Merisiers, Gif-sur-Yvette 91191, France

Ivison RJ. Ibar E. : Royal Observatory, UK Astronomy Technology Centre, Blackford Hill, Edinburgh EH9 3HJ, UK

Michalowski M. : Institute for Astronomy, University of Edinburgh, Blackford Hill, Edinburgh EH9 3HJ, UK

Negrello M. Hopwood R. : Department of Physics and Astronomy, Open University, Milton Keynes MK7 6AA, UK

Aretxaga I. Hughes DH. : Instituto Nacional de Astrofísica, Óptica y Electrónica, Aptdo. Postal 51 y 216, Puebla 72000, Mexico

Auld R. Dariush A. Eales SA. Pascale E. Pohlen M. : School of Physics and Astronomy, Cardiff University, The Parade, Cardiff CF24 3AA, UK

Baes M. Fritz J. : Sterrenkundig Observatorium, Universiteit Gent, Krijgslaan 281, S9, Ghent B-9000, Belgium

Buttiglione S. Jarvis M. de Zotti G. : INAF-Osservatorio Astronomico di Padova, Vicolo dell'Osservatorio 5, Padova I-35122, Italy

Cava A. : Departamento de Astrofísica, Universidad Complutense de Madrid, Madrid E-28040, Spain

Cooray A. : Department of Physics and Astronomy, University of California, Irvine, CA 92697, USA

Dunne L. Dye S. Maddox S. Rigby E. Smith DJB. : School of Physics and Astronomy, University of Nottingham, University Park, Nottingham NG7 2RD, UK

Frayser D. : National Radio Astronomy Observatory, P.O. Box 2, Green Bank, WV 24944, USA

Swinbank AM. : Institute for Computational Cosmology, Durham University, South Road, Durham DH1 3EE, UK

Temi P. : Ames Research Center, Astrophysics Branch, NASA, Mail Stop 245-6, Moffett Field, CA 94035, USA

Valtchanov I. : Herschel Science Center, ESA, P.O. Box 78, Villanueva de la Canada, Madrid 28691, Spain

van der Werf P. : Leiden Observatory, Leiden University, P.O. Box 9513, Leiden 2300 RA, Netherlands

Title

Gas and Dust in a Submillimeter Galaxy at $z = 4.24$ from the Herschel Atlas

Source

Astrophysical Journal, vol.740, no.2, 20 Oct. 2011, 63 (10 pp.). Publisher: IOP Publishing Ltd., UK.

Abstract

We report ground-based follow-up observations of the exceptional source, ID 141, one of the brightest sources detected so far in the Herschel Astrophysical Terahertz Large Area Survey cosmological survey. ID 141 was observed using the IRAM 30 m telescope and Plateau de Bure interferometer (PdBI), the Submillimeter Array, and the Atacama Pathfinder Experiment submillimeter telescope to measure the dust continuum and emission lines of the main isotope of carbon monoxide and carbon ([C I] and [C II]). The detection of strong CO emission lines with the PdBI confirms that ID 141 is at high redshift ($z = 4.243 \pm 0.001$). The strength of the continuum and emission lines suggests that ID 141 is gravitationally lensed. The width ($V_{\text{FWHM}} \sim 800 \text{ km s}^{-1}$) and asymmetric profiles of the CO and carbon lines indicate orbital motion in a disk or a merger. The properties derived for ID 141 are compatible with an ultraluminous ($L_{\text{FIR}} \sim (8.5 \pm 0.3) \times 10^{13} L_{\odot}$, where L is the amplification factor), dense ($n \sim 10^4 \text{ cm}^{-3}$), and warm ($T_{\text{kin}} \sim 40 \text{ K}$) starburst galaxy, with an estimated star formation rate of $(0.7-1.7) \times 10^4 L_{\odot} \text{ yr}^{-1}$. The carbon emission lines indicate a dense ($n \sim 10^4 \text{ cm}^{-3}$) photon-dominated region, illuminated by a far-UV radiation field a few thousand times more intense than that in our Galaxy. In conclusion, the physical properties of the high- z galaxy ID 141 are remarkably similar to those of local ultraluminous infrared galaxies. (69 References).